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Magnetic fabric of the Parashi stock and related dyke swarm, Alta Guajira (Colombia): The Caribbean-South American plates oblique convergence

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ABSTRACT. Anisotropy of magnetic susceptibility (AMS) and anhysteretic remanence (AAR) were used to evaluate the emplacement history of the Parashi stock and related dyke swarm situated in NW Colombia. The average magnetic susceptibility of 4.5×10^{-2} SI, in conjunction with low-coercivity components provided by the isothermal remanence and thermomagnetic curves with net Verwey and Curie transitions, indicates that multidomain magnetite records the anisotropy directions. The similar orientation and shape of the AMS and AAR ellipsoids indicate the absence of very fine magnetite with an inverse fabric. The magnetic foliation is the best-defined fabric element in these rocks and outlines a concentric structure, elongated parallel to the NE-SW direction of the pluton. Crystallisation age of the stock and dykes (51-47 Ma), along with pressure of emplacement determination indicate that the stock and the dyke swarm probably formed simultaneously, and they were emplaced in the shallow crust (<10 km). The magnetic fabric agrees with emplacement in a transtensional setting, with the dyke swarm infilling tensional cracks formed during the final stages of pluton crystallisation. This tectonic setting likely records the difficulty of the thick oceanic crust subducting, which would result in general transtension in the overriding continental plate and cessation of arc magmatism in the Eocene.

Keywords: AMS, Petrofabric, Pluton emplacement, Parashi stock, Caribbean plate.

RESUMEN. Fábrica magnética del stock de Parashi y el enjambre de diques relacionado, Alta Guajira (Colombia): la convergencia oblicua entre las placas del Caribe y Sudamericana. Se utilizó Anisotropía de Susceptibilidad Magnética (ASM) y Remanencia Anisterética (ARA) para evaluar la historia de emplazamiento del stock Parashi y del enjambre de diques relacionados situados en el noroeste de Colombia. La susceptibilidad magnética promedio de 4,5×10⁻² SI, junto con los componentes de baja coercitividad proporcionados por la remanencia isotérmica y las curvas termomagnéticas con las transiciones de Verwey y Curie, indican que la magnetita multidominio registra las direcciones de anisotropía. Las orientaciones y formas similares de los elipsoides ASM y ARA sugieren la ausencia de magnetita muy fina con una fábrica inversa. La foliación magnética es el elemento de la fábrica mejor definido en estas rocas y determina una estructura concéntrica, alargada paralela a la dirección NE-SO del plutón. La edad de cristalización del stock y los diques (51-47 Ma), junto con los cálculos de presión de emplazamiento, indican que el stock y el enjambre de diques probablemente se formaron simultáneamente y se emplazaron en la corteza poco profunda (<10 km). La fábrica magnética concuerda con el emplazamiento en un régimen transtensional y con el enjambre de diques que rellena las fracturas tensionales formadas durante las etapas finales de la cristalización del plutón. Este escenario tectónico probablemente registra la dificultad de subducción de la corteza oceánica gruesa, lo que resultaría en una transtensión general en la placa continental y el cese del magmatismo de arco en el Eoceno.

Palabras clave: ASM, Petrofábrica, Emplazamiento de plutones, Stock Parashi, Placa del Caribe.

1. Introduction

Low-field Anisotropy of Magnetic Susceptibility (AMS) is a useful technique for recording fabrics in rocks that appear isotropic by visual inspection. This anisotropy is based on the difference between the magnetic susceptibility measured in different directions of the rock, which itself depends on the preferred orientation of all the mineral grains weighted according to their concentrations and grain anisotropies (Hrouda, 1982). In granitic rocks, AMS records the bulk contribution of the diamagnetic (quartz, feldspars), paramagnetic (micas, amphibole) and ferromagnetic (magnetite, hematite) mineral assemblages. The relationship between AMS and magmatic fabrics in plutons, irrespective of the paramagnetic and/or ferromagnetic source of the AMS, has been widely used to investigate emplacement mechanisms and relationships with regional tectonic processes (Archanjo et al., 1994; Cruden, 1998; Siegesmund and Becker, 2000; Clemens and Benn, 2010; Trubac et al., 2014; Mamtani et al., 2019).

The Anisotropy of Anhysteretic Remanence (AAR) is a complementary magnetic method that records only the contribution to the magnetic fabric of minerals carrying a remanence, which is mostly magnetite in granitic rocks (Jackson, 1991). Like AMS, AAR can also be represented as a second-rank tensor, with the principal directions represented by an ellipsoid. The AMS and AAR data, therefore, can provide valuable information about emplacement tectonics at different stages of pluton construction (Bouchez, 1997), especially in volcano-plutonic complexes in which earlier intrusions are often later invaded, usually in the shallower crust, by subvolcanic rocks and dykes (*e.g.*, Magee *et al.*, 2012; Tomek *et al.*, 2017; Burton-Johnson *et al.*, 2019).

In this contribution, AMS and AAR fabric data have been used to evaluate the emplacement mechanisms of a shallow quartz diorite stock, intruded by andesitic to dacitic dykes, situated in the Guajira Península of northern Colombia. A Late Cretaceous to Eocene magmatic arc is present along the Andes, from Ecuador to the Caribbean margin of Colombia and Venezuela. This is related to the oblique subduction of the Caribbean oceanic plate beneath the northern edge of the South American plate (Fig. 1; Pindell et al., 2005; Hughes and Pilatasig, 2002; Vallejo et al., 2006; Pindell and Kennan, 2009; Bustamante et al., 2009; Cardona et al., 2011a, b; Weber et al., 2010;

Montes *et al.*, 2019). The Parashi stock and dykes intrude Cretaceous accretionary complexes associated with a former arc-continent collision (Cardona *et al.*, 2009; Weber *et al.*, 2009, 2010; Montes *et al.*, 2019) and constitute one of the latest records of Eocene magmatism in Colombia (Cardona *et al.*, 2014). This study examines the magnetic fabric of a quartz-dioritic stock and related dyke swarm from a time when the oblique convergence of the Caribbean plate triggered the exhumation and disruption of the former Late Cretaceous collisional orogen in the Guajira region.

2. Geologic setting

The Late Cretaceous to Paleogene tectonic evolution of northwestern Andean Cordillera and northern South American-Caribbean region has been related to the interaction between the Caribbean plate, as it migrated from a southwestern Pacific position, and the continental margin of South America (Burke et al., 1984; Lugo and Mann, 1995; Kerr et al., 1997; Pindell et al., 1998, 2005; Pindell and Keenan, 2009; Boschman et al., 2014; Montes et al., 2019). The continuous oblique convergence between the two tectonic plates is responsible for major block displacements and the formation of Cenozoic basins in northernmost Colombia and Venezuela (Duque-Caro, 1979; Muessig, 1984; Macellari, 1995; Montes et al., 2010). The subsequent subduction of the Caribbean plate beneath the South American continental margin is responsible for the formation of a short-lived Paleocene-Eocene magmatic arc (Cardona et al., 2011b, 2014; Bayona et al., 2012; Bustamante et al., 2017). The oblique nature of the Caribbean-South American plates interactions has been inferred mainly from palaeogeographic reconstructions, including palaeomagnetic evidences of block rotation, lithotectonic correlation and basin filling histories (Escalona and Mann, 2003; Pindell and Keenan, 2009; Montes et al., 2010, 2019; Boschman et al., 2014; Nova et al., 2019).

The Palaeogene magmatic arcs formed in continental and oceanic domains that resulted from the Caribbean convergence, are exposed in the Central and Western Cordilleras of Colombia (Cardona *et al.*, 2018). In the Central Cordillera, the magmatism lasted from ~60 to 45 Ma, and the related plutonic rocks intruded a thickened continental margin that grew as a consequence of the former Caribbean-

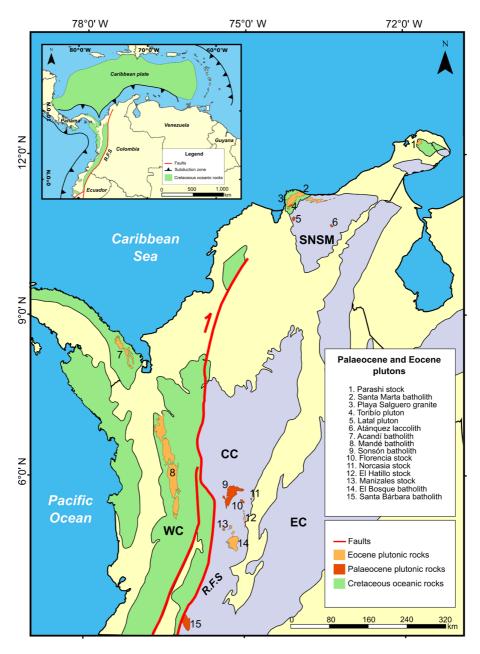


FIG. 1. Regional map showing the current tectonic context of the Caribbean (inset). The subduction zones of the Antilles (east) and Panama (west), and strike-slip faults to the north and south, are limits of this plate. Accreted oceanic terranes related to the oblique subduction of the Caribbean plate beneath northwestern South America crops out along western Colombia, Ecuador and Panama, to the west of the Romeral Fault System (R.F.S). The larger map is showing the Cretaceous oceanic rocks of the West Cordillera (WC) in contact with the Central Cordillera (CC) along the Romeral Fault System (R.F.S). Plutonic rocks indicated with numbers corresponds to the Palaeocene-Eocene magmatic arc. EC: Eastern Cordillera; SNSM: Sierra Nevada de Santa Marta (modified from Bayona et al., 2012; Kerr, 2014 and Salazar et al., 2016).

South American plate collision (Bayona *et al.*, 2013; Bustamante *et al.*, 2017; Cardona *et al.*, 2018; Duque-Trujillo *et al.*, 2019a). In the Western Cordillera, the

volcanic and plutonic complexes formed from 71 Ma to 34 Ma and probably intrude remnants of an Early Cretaceous oceanic plateau (Cardona *et al.*, 2018).

Farther north in the Caribbean region (Sierra Nevada de Santa Marta, Guajira Peninsula), the magmatic rocks intruded subduction- and/or collisional-related complexes (Weber et al., 2009, 2010; Cardona et al., 2010). One of these collisional complexes is the Late Cretaceous Etpana Formation that crops out on the northwestern part of the Guajira Península (Fig. 2) and comprises low-grade greenschists, phyllites and quartzites, and some intercalated serpentinites, rodingites and gabbros (Lockwood, 1965; Weber et al., 2010). Lockwood (1965) describes the foliation in the Etpana Formation as defined by a weak cleavage that is usually oriented parallel to bedding and with nearly E-W trending folds developed in the southern portions of the formation and superimposed on N-S folds developed north and east of the Parashi stock (Fig. 2).

Among the most studied Paleogene plutons of the Caribbean region are the Santa Marta batholith and the

Parashi stock (Cardona *et al.*, 2011b, 2014; Salazar *et al.*, 2016; Duque-Trujillo *et al.*, 2019b). The former corresponds to the largest Eocene pluton in Colombia, whereas the Parashi stock and sub-volcanic rocks found in well-logs near the Paraguana Península in Venezuela mark the northeastern end of the Eocene magmatism (Bayona *et al.*, 2012; Bustamante *et al.*, 2017; Baquero *et al.*, 2014). The absence of post-Eocene magmatic rocks is likely due to the difficulty of the thick oceanic plateau to subduct combined with the more oblique nature of the Caribbean-South American plates convergence (Bayona *et al.*, 2012; Cardona *et al.*, 2014; Montes *et al.*, 2019).

The Santa Marta batholith consists of granodiorites and tonalites, dated between 58 and 54 Ma (U-Pb in zircons), which are followed by small volume of felsic pulses emplaced between 52 and 50 Ma (Cardona *et al.*, 2011b; Salazar *et al.*, 2016; Duque-Trujillo *et al.*, 2019b). The AMS and AAR results

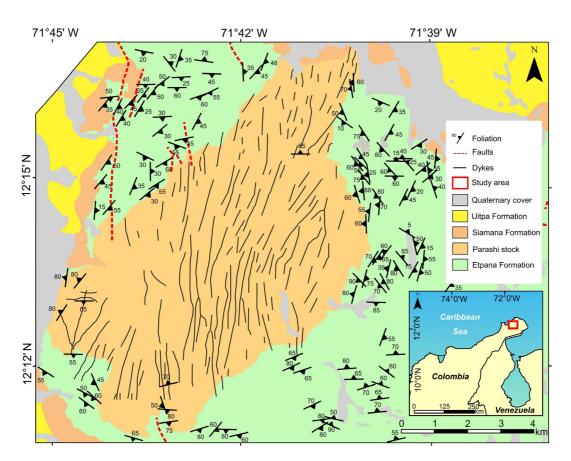


FIG. 2. Geological map showing the location of the Parashi stock, the trend of the dyke swarm and the foliations of the country rock (After Lockwood, 1965).

from the Santa Marta batholith have been interpreted to record a transpressive arc setting related to the compression and lateral migration of the Caribbean plate (Salazar et al., 2016). The Parashi stock (about 56 km²; Fig. 2), in turn, is located to the northeast of the Santa Marta batholith in the Jarara Serranía of the Guajira Península. It is surrounded by middle Tertiary marine limestones and shales of the Uitpa and Siamana formations (Fig. 2) and consists of a quartz diorite to granodiorite body with biotite and hornblende that is partially covered unconformably by Oligo-Miocene conglomerate and carbonate rocks (Lockwood, 1965; Zapata et al., 2010). The stock includes a swarm of N- to NE-trending andesitic to dacitic dykes, and all these rocks intrude lowgrade metapelites and the metaultramafic rocks of the Etpana Formation (Fig. 2). A thermal aureole of approximately 50 m in width has been described in the metapelites around the Parashi stock. These rocks reached the hornblende-hornfels facies, as attested to by the presence of andalusite crystals set in a matrix of quartz, biotite and oligoclase, which is shown to be replacing a former lower-temperature assemblage containing sericite, chlorite, quartz and albite (Lockwood, 1965; Martínez and Zuluaga, 2010; Weber et al., 2010). Both the pluton and the dykes yielded zircon U-Pb crystallisation ages between 51 and 47 Ma which, in turn, overlaps with the late crystallisation phases of the Santa Marta Batholith (Cardona et al., 2014).

3. Methods

3.1. Sampling and sample preparation

We collected samples from 13 sites in the Parashi stock and 11 sites from different dykes. The relatively few samples collected in the pluton and dykes are related to poor exposure of the rocks in the Guajira region as well as the limited access due to the few roads available. For the magnetic studies, we drilled 2 or 3 oriented rock cores of one-inch diameter using a portable gasoline-powered drill. In the laboratory, the cores were sliced into 2.2 cm long specimens, providing between 5 and 8 specimens per site in both pluton and dykes. A total of 84 specimens for the stock and 67 specimens for the dykes were studied by both AMS and AAR methods in the *Laboratório de Paleomagnetismo* at the University of São Paulo, Brazil.

3.2. Anisotropy of magnetic susceptibility

Low-field AMS was measured on an MFK1-FA Kappabridge (200 A/m, AC field at 920 Hz), and the AMS tensor was calculated using Jelinek statistics (Jelinek, 1978) through ANISOFT software provided by the equipment manufacturer (AGICO). The calculated AMS parameters are the principal directions of the ellipsoid tensor $(k_1 \ge k_2 \ge k_3)$, the mean susceptibility $(k_m = (k_1 + k_2 + k_3)/3)$, the anisotropy degree $(P = k_1/k_3)$ and the ellipsoid shape $(T = [2\ln(k_1/k_1)/(k_2/k_3)] - 1)$, such that $-1 \le T < 0$ for prolate and $0 < T \le 1$ for oblate ellipsoids. The AMS results are shown in table 1.

3.3. Anisotropy of anhysteretic remanence

The AAR technique allows to analyse the (ferro)magnetic phases that contribute to the remanence. The method consists of inducing an anhysteretic magnetization in different orientations of the specimen and measuring the respective remanence in each orientation (see details in McCabe et al., 1985 and Jackson, 1991). The resulting anisotropy of anhysteretic remanence tensor can be described by an ellipsoid and then compared to the AMS results. This approach is particularly useful in detecting sub-fabrics (Jackson, 1991; Borradaile and Jackson, 2004) and the eventual contribution of very fine-seized magnetic particles in the rock that may provide inverse magnetic fabrics (Stephenson et al., 1986; Rochette et al., 1992). The magnitude, shape and principal directions of the AAR tensor were determined with ANISOFT software.

Coercivity allows estimating the size of the ferromagnetic minerals due to the dependence of the Anhysteretic Remanent Magnetization (ARM) on the grain size (Jackson et al., 1988). The coercivity spectrum is determined by demagnetising the specimen in a peak AC field of 100 mT, whereas a steady anhysteretic DC field is applied between two chosen intervals of a decaying AF peak (see details in Jackson et al., 1988). The alternating field coupled with a steady anhysteretic direct field was applied by a demagnetizer (LDA) from AGICO and the remanent magnetization measured in a JR5 magnetometer. The coercivity spectra were studied at five sites of the stock and five for the dykes.

TABLE 1. AMS PARAMETERS OF THE PARASHI STOCK AND DYKES.

Site	Location*	Km	P	T	K1dec	K1inc	K2dec	K2inc	K3dec	K3inc
				F	Pluton					
CP11	N12 15 22.8 W71 41 54.8	52,712	1.12	-0.02	83	13	260	76	276	3
CP12-2	N12 16 48.2 W71 40 33.3	50,576	1.07	0.67	200	29	223	46	167	24
CP13	N12 16 32.2 W71 40 27.9	52,425	1.11	0.44	183	79	175	6	162	9
CP14	N12 15 58.1 W71 40 30.3	56,383	1.10	0.59	159	30	255	59	129	4
CP15	N12 15 24.8 W71 40 49.9	43,787	1.18	0.61	287	14	71	62	218	21
CP16	N12 15 03.9 W71 40 50.6	49,093	1.19	0.68	292	49	103	40	130	7
CP20	N12 16 16.7 W71 41 13.5	56,981	1.11	0.47	217	37	95	52	192	5
CP22	N12 15 00.2 W71 41 24.6	47,095	1.17	0.53	89	55	309	28	208	18
CP24	N12 13 40.9 W71 43 21.6	25,069	1.20	0.47	138	14	284	57	197	29
CP25	N12 13 24.1 W71 44 07.6	49,232	1.17	0.73	59	56	211	31	306	9
CP26	N12 13 54.0 W71 43 39.2	31,151	1.18	0.67	87	37	194	48	310	11
CP7	N12 13 33.1 W71 40 45.6	30,289	1.15	0.53	121	46	101	17	287	37
CP8	N12 13 36.8 W71 40 32.7	30,612	1.12	0.68	38	42	161	33	278	28
CP9	N12 14 28.2 W71 40 15.0	40,147	1.20	0.63	242	6	224	15	44	74
				1	Dykes					
CP1	N12 15 18.1 W71 42 31.8	8,371	1.02	-0.07	341	30	145	58	217	7
CP10	N12 14 47.8 W71 40 51.8	10,242	1.06	0.03	242	42	49	47	171	8
CP12-1	N12 16 48.2 W71 40 33.3	22,569	1.04	-0.52	211	32	220	27	168	43
CP17	N12 13 41.2 W71 41 54.3	3,025	1.03	0.56	180	38	70	49	308	12
CP18	N12 14 36.0 W71 41 01.4	7,219	1.03	-0.04	235	22	340	35	119	47
CP19	N12 16 50.9 W71 41 06.0	34,161	1.04	0.15	131	49	121	22	248	25
CP21	N12 15 32.5 W71 41 24.8	13,420	1.02	0.01	98	48	58	11	266	40
CP23	N12 13 53.2 W71 43 06.1	11,979	1.09	0.34	165	9	221	69	21	17
CP3	N12 14 49.6 W71 42 02.5	7,820	1.05	0.29	295	17	192	38	44	46
CP4	N12 14 21.8 W71 42 10.3	4,971	1.03	-0.06	51	24	234	56	190	19
CP5	N12 14 12.2 W71 41 02.9	14,529	1.03	0.04	252	55	76	16	105	30
CP6	N12 13 50.7 W71 40 53.9	5,506	1.04	0.26	73	17	195	69	194	12

^{*}Bogotá coordinate system.

km: mean susceptibility; P: degree of anisotropy; T: shape of the AMS ellipsoid; K1 and K3: magnetic lineation and foliation (inclination: inc and declination: dec) respectively.

3.4. Magnetic mineralogy

The magnetic mineralogy was investigated through the temperature dependence of the magnetic susceptibility (k-T). Representative samples of the stock were crushed to a fine sand size and the susceptibility of a small volume fraction measured from -200 to 700 °C using a CS-3/CSL

furnace attached to the KLY-4 Kappabridge. At low temperatures, the susceptibility was recorded by warming the sample from -200 °C to room temperature. From room temperature up to 700 °C (heating cycle) and back (cooling cycle), k-T was measured in an argon environment to prevent the oxidation of the sample from atmospheric oxygen during the heating/cooling cycle.

4. Results

4.1. Rock textures and microstructures

Thin sections of every AMS site were analysed in order to identify the mineral textures that could be related to the magnetic information. Petrographic analyses performed on the Parashi stock (Lockwood, 1965; Cardona *et al.*, 2014) show that the pluton is relatively homogeneous in composition, varying dominantly from quartz diorite to, locally, granodiorite. The texture in thin-section is typically allotriomorphic, where plagioclase and quartz compose ~70%-80% and K-feldspar commonly less than 5% of the rock. Ferromagnesian minerals consist mostly of homblende and biotite, and these make up ~20% of the rock. Accessory minerals include zircon, apatite, sphene, allanite and opaque minerals.

Although characterised by medium- to coarse-grained typically magmatic textures, the microstructures of the Parashi stock record widespread internal deformation, such as elongate recrystallised aggregates of quartz, deformation twins in plagioclase, the development of myrmekite and microcline twinning. Furthermore, chessboard quartz crystals often with highly sutured grain boundaries evidence that the deformation occured at high temperature (Vernon, 2000). As the host metabasic rocks usually show low-grade metamorphic conditions (Weber *et al.*, 2010), we consider that the internal deformation of the pluton mineral assemblage was acquired at the latest stages of magmatic crystallisation and/or just after its full crystallisation but still at high temperature.

Dykes are mainly dacite-phyric in composition and trend to N- to NE-direction. Euhedral phenocrysts of plagioclase and quartz are settled in a finer quartz-feldspathic matrix that includes biotite and hornblende. Ferromagnesian minerals show a preferred orientation, visible in sections parallel to the long axis of the amphiboles and/or perpendicular to the basal plane of biotite grains (Fig. 3). The mineral lineation defined by the preferred orientation of the amphiboles (320°, 40°) are usually consistent with the orientation of the magnetic lineation.

4.2. Susceptibility and magnetic mineralogy

Mean susceptibility (k_m) values of the Parashi stock are very high, ranging from 24×10^{-3} to 64×10^{-3} SI, with an average of 45×10^{-3} SI (45 mSI), whereas

the mean anisotropy degree (P) varies from 1.09 to 1.22 (Fig. 4). The dykes also show high k_m values, ranging from 2.11 to 37.8 mSI (with a mean of 11.6 mSI), and the dykes degree of anisotropy ranges from 1.01 to 1.10 (Fig. 4). The AMS ellipsoids are mainly oblate (Fig. 4; 91% of the specimens have T > 0.2), though a few specimens display a neutral shape (Fig. 4; -0.2<T<0.2 in 8% of the specimens). The dykes have ellipsoids that vary from oblate to prolate (Fig. 4), although most of the ellipsoids (47%) show neutral shapes.

The k-T curves of the Parashi stock display a net Verwey transition around -150 °C and a strong drop of susceptibility on heating close to 580 °C (Fig. 5). On cooling, the curve is nearly reversible, indicating that none, minor new phases were formed in the experiment during the heating cycle. The coercivity spectra of the granitic rocks are similar to each other (Fig. 5) with a peak of coercivity at 10 mT that practically vanishes after 30 mT. These results, combined with the k-T curves, indicate that a coarse, low coercive, Ti-poor magnetite dominates the susceptibility and anisotropy of the Parashi stock.

The k-T curves of the dykes show the distinct Verwey transition and Curie temperature of Tipoor magnetite, although some samples record some differences compared to the pluton. During the heating cycle between ~150 and 250 °C, the curve displays a sinuous appearance (Fig. 5), which suggests a phase transition between 200 and 300 °C of and oxidised magnetite (maghemite) and/ or hematite (Gehring et al., 2009) as evidenced, in addition, by the irreversibility of the susceptibility values upon cooling. However, hematite, which has a Morin transition at -10 °C and a Néel temperature at 670 °C, was not recorded in the k-T experiments. The coercivity spectrum of the dyke samples also decrease after a remanence peak at 10 mT. However, for samples with the lowest susceptibilities, the remanence decreases progressively and is not fully removed at 100 mT (Fig. 5). This indicates that these samples contain, in addition to low coercive magnetite, some high coercive fraction probably corresponding to fine-grained magnetite and hematite.

4.3. Magnetic fabric

The k-T and coercivity spectra indicate that coarse Ti-poor magnetite carries the magnetic anisotropy of the pluton and dykes. Accordingly,

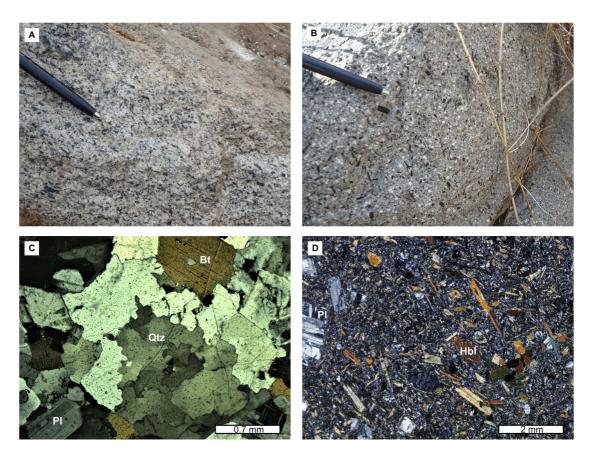


FIG. 3. Representative macrostructures of the (A) Parashi stock, showing oriented amphibole crystals that define the foliation; and (B) dykes with plagioclase and tabular amphibole phenocrysts in a quartz-feldspathic matrix. Microstructures in the (C) Parashi stock shows sutured boundaries and undulatory extinction in quartz, and kink bands in biotite; and (D) the dykes show subtle aligned hornblende crystals in a microcrystalline matrix.

although the AAR is slightly higher than the AMS in both the stock and the dykes, AMS and AAR scalar parameters are similar to each other, including their respective principal directions (Fig. 6).

Some differences, however, are recorded in the AMS parameters of the stock and dykes, such as the higher mean susceptibility and degree of anisotropy of the pluton (Fig. 6). In addition, the AMS ellipsoid shapes are dominantly oblate in the pluton and vary from oblate to prolate in the dykes (Fig. 4). Therefore, the magnetic foliation is the best-defined fabric element recorded in the pluton.

The AMS and AAR foliations (defined as perpendicular to the k_3 axes) in the pluton tend to be parallel to the west and east contacts with the wall rocks and highly oblique to perpendicular to the pluton contacts in its northern part (Fig. 7). If the

magnetic foliations and the planar structures measured in the field (Fig. 7) are combined, a net concentric foliation with moderate to steep dips emerges for the pluton. Magnetic lineation (k_1) , in turn, show a more organised NE trend with a moderate to steep plunge, although some E-trending lineation are evident in the northern part of the intrusion (Fig. 7).

Dykes yield fabrics that can be classified as normal type, *i.e.*, with magnetic foliation parallel to slightly oblique to the dyke margin (Rochette, 1987). However, two dykes (CP3 and CP23) show an inverse-type fabric in that both the foliation and lineation are nearly perpendicular to the dyke plane (Fig. 8). Nevertheless, as the respective AMS and AAR principal directions are coaxial (Fig. 6), the presence of very fine, single-domain magnetite that produce anomalous fabric types can be discarded (Rochette, 1987).

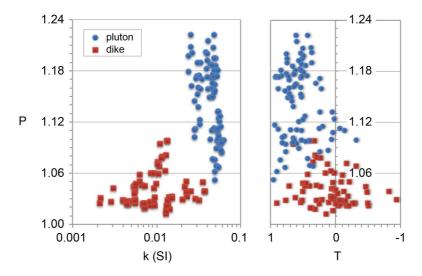


FIG. 4. Relation between the mean site magnetic susceptibility (k), the degree of anisotropy (P) and the shape (T) of the AMS ellipsoid.

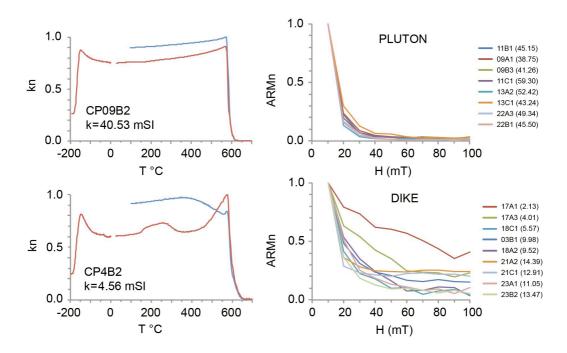


FIG. 5. **Left side:** dependence of magnetic susceptibility (kn) with temperature (T °C) for representative samples of the Parashi stock (CP09B2) and related dykes (CP4B2). **Right side:** coercivity spectra for samples with different susceptibility magnitudes. H induced field, *ARMn* normalised anhysteretic remanent magnetisation.

5. Discussion

The distinct Verwey transition with a net drop of susceptibility at 580 °C upon heating (Fig. 5), and the dominance of low-coercivity fractions on the

ARM spectra (Fig. 5) indicate that multidomain, coarse magnetite controls the anisotropy of magnetic susceptibility in the Parashi stock and the associated dykes (Muxworthy and McClelland, 2000; Trindade *et al.*, 2001). Although a few dyke samples displayed

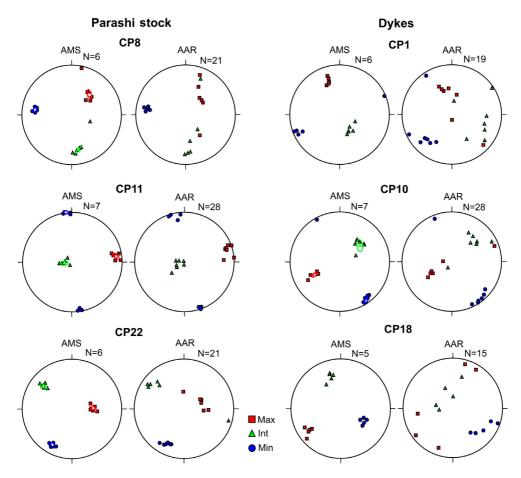


FIG. 6. Stereograms (lower hemisphere) of the principal directions of AMS and AAR for selected samples from the Parashi stock and dykes.

lower susceptibilities and k-T curves that suggest minor contributions of possibly hematite and/or maghemite, the consistency between the AMS and AAR indicates these phases do not modify the magnetic fabric orientation. Late magmatic-hydrothermal fluids associated with the emplacement of the dykes might account for the crystallisation of the secondary magnetic phases, such as hematite which, in some cases, can partially replace the coarse magnetite. Fine magnetite particles, which could have been formed by hydrothermal fluids and hence might provide anomalous fabrics, are precluded due to the similarity between the AMS and AAR principal directions (Fig. 6; Rochette et al., 1992). Magnetic fabrics recorded in the Parashi magmatic complex are therefore considered to depend on the preferred orientation of the magnetite grains which, according

to comparative studies with silicate fabrics in granitic rocks, closely approximate the magmatic fabric fairly well (Archanjo *et al.*, 1995; Salazar *et al.*, 2008; Arbaret *et al.*, 2013; Schopa *et al.*, 2015).

Foliations defined by the AMS data, combined with the planar structures measured within the pluton, are consistent with a concentric planar fabric acquired during the emplacement, possibly resulting from the action of internal body (ascensional) forces. A similar fabric type was recorded in experimental models simulating the upward extrusion of lava domes (Harnett *et al.*, 2018). The concentric fabric would be formed by drag forces between the wall rocks and a confined volume of magma moving upward, perhaps being enhanced by helical magma movements that arose when the flow velocity in the inner part of the conduit was higher than in its outer part (Fowler, 1996;

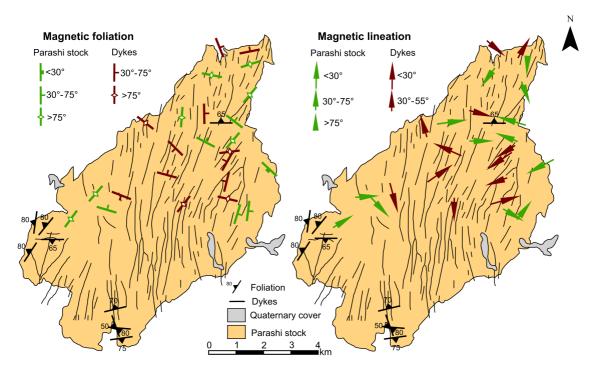


FIG. 7. Magnetic fabric (foliation and lineation) showing a map with the foliation (**left**) and lineation (**right**) of samples from the Parashi stock and dykes. Field structural data (in black) are also included within the pluton's map.

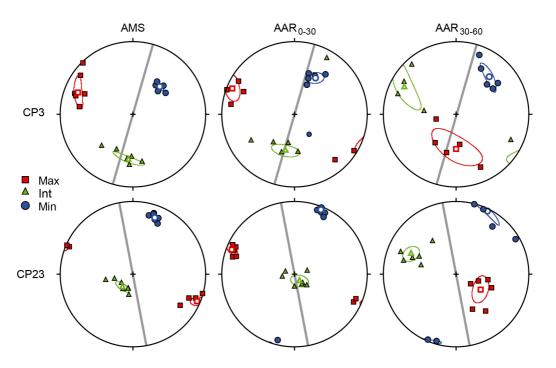


FIG. 8. AMS and partial AAR (0-30; 30-60) for dykes showing an inverse-type fabric (CP3 and CP23), where the grey line represents the dyke trend measured in the field.

Trubac *et al.*, 2009). The differential displacement of the magma would additionally be able to deform the magmatic microstructures at high temperature, as shown by the typical fabrics of the Parashi stock.

The magnetic fabrics of the dykes are typically of normal type, *i.e.*, formed during the crystallisation and cooling of the magma confined between the dyke walls. However, two dykes show an anomalous magnetic fabric (CP3 and CP23) in that the lineation is perpendicular to the walls. As the AAR principal directions are coaxial to the AMS (Fig. 6), the presence of very fine magnetite particles with an inverse fabric is discarded. We can speculate, therefore, that the magnetic lineation would record the direction of the brittle opening of the dyke wall, with the magnetite crystals growing parallel to the local extension direction of the dyke.

Evidence for the shallow emplacement of the stock is attested by the intrusion of sub-volcanic felsic dyke swarm and wall rock structures that are partially truncated at the stock margin. Pressure calculations for the Parashi stock were estimated using the available whole-rock geochemistry (Cardona et al., 2014) and the Yang (2017) geobarometer. This method is based on the inverse relationship between SiO₂ content and pressure of emplacement. Based on geochemical data reported by Cardona et al. (2014), samples with SiO₂> 60 wt% were chosen. Pressures were obtained after the normative composition was calculated, using the values of quartz (Qz), albite (Ab), and orthoclase (Or). The emplacement depth was calculated from the pressure values, taking into account the relationship of 100 MPa=2.7 km (Yang, 2017). The average pressure results (Table 2) of the Parashi stock are 290 MPa, corresponding to an emplacement depth of ca. 8 km. This estimate is similar to those reported by Cardona et al. (2011a), who used the Al in hornblende barometer of Schmidt (1992) and obtained pressures ranging from 230 MPa to 300 MPa, or 7 km to 9 km of emplacement depth. These barometric data indicate that the Parashi stock emplaced in the shallow crust and that, in the latest stages of crystallisation or after fully crystallised, the stock deformed brittlely and facilitated the emplacement of the dyke, with the magma likely extracted from a reservoir that remained in the deeper crust. The shallow emplacement of the pluton is also attested by the local and alusite hornfels formed next to the contact between the pluton and the low-grade Etpana schist.

TABLE 2. EMPLACEMENT PRESSURE CALCULATIONS.

Dept (km		10	9	6	∞	9
) Pressure Average (Kbar)		4	2	3	3	2
	Average	385	238	331	308	227
Pressure (Mpa)	Equation2 (Ab+Or)	380	233	326	303	223
Pre	Equation1 (Q)	390	243	336	312	232
	Ab	42.0	38.4	32.8	32.1	32.5
Norm	Or	13.6	11.4	11.7	99.8 19.1 11.0 32.1	12.1
	ò	22.9	23.8	19.2	19.1	21.6
	Total	99.6 22.9 13.6 42.0	7.66	7.66	8.66	2.0 0.2 0.5 99.7 21.6 12.1 32.5
	T01	2.2	0.4	0.7	8.0	0.5
	P_2O_5	0.1 2.2	0.1	0.2	0.2	0.2
	K_2^{\dagger}	2.3	1.9	2.0	1.9	2.0
	Na_2O	4.8	4.5	3.8	3.8	
	CaO	2.9	4.1	5.5	5.7	5.1
	$_{ m MgO}$	0.1 1.5 2.9 4.8	1.3	2.0	2.1	4.7 0.1 1.7 5.1 3.8
	MnO	0.1	0.1	0.1	0.1	0.1
	$\mathrm{Fe}_{\scriptscriptstyle 2}\mathrm{O}_{\scriptscriptstyle 3}$	2.4	2.8	5.2	5.2	4.7
	Al_2O_3	16.0	16.8	17.2	17.4	16.8
	SiO ₂ TiO ₂ Al ₂ O ₃ Fe ₂ O ₃ MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ LOI Total Q Or Ab	0.3	0.3	9.0	9.0	0.5
	SiO_2	67.2	67.5	62.5	62.1	64.1
		Dyke	Dyke 67.5	Stock 62.5	Stock 62.1 0.6	Stock
	Sample	CM3-11p Dyke 67.2 0.3 16.0	CM3-7p	CM5-19	CM5-21	CM5-22g Stock 64.1 0.5 16.8

Equations 1 [P=0.2426(Q)³+26.392(Q)²-980.74(Q)+12,563] and 2 [P=0.2426(Ab+Or)³ 46.397(Ab+Or)²+2,981.3(Ab+Or)-64,224] are from Yang (2017). Chemical composition of the stock and dykes in wt%. Kbar pressure is rounded off. Q: quartz, Or: orthoclase, Ab: albite.

Different palaeogeographic reconstructions, based on regional and local field relations, suggest that the interaction between the Caribbean plate and northern South American plate, since the Late Cretaceous to Palaeocene, has been dominated by oblique convergence and is characterised by continuous changes in the convergence relationship (Pindell et al., 1998; Müeller et al., 1999; Montes et al., 2005, 2010, 2019; Pindell and Keenan, 2009). In the Colombian margin, such interactions began with a Late Cretaceous-Palaeocene arc-continent collisional event, followed by the inception of a new subduction zone after 63 Ma, as a consequence of the subduction of the Caribbean plate beneath the northern edge of the South American plate. The subduction appears to have continued until 45 Ma, when it switched off, potentially as a consequence of an increase of plate convergence obliquity (Cardona et al., 2010, 2011b; Bayona et al., 2012; Bustamante et al., 2017; Montes et al., 2019). This setting may have created the room needed for the Parashi stock to be emplaced into the shallow crust, where it acquired its concentric fabric, due to the development of the transtensional regime.

According to Cardona *et al.* (2014), the crystallisation age of the Parashi stock and the dykes is ~51 to 47 Ma, and this time interval overlaps in analytical error. It, therefore, does not permit a clear distinction of the duration of each crystallisation event, although it seems clear that there was a very short time interval between both events based on these ages. As has been described in similar tectonic settings where strong oblique convergence

dominated the deformation partitioning (Glazner *et al.*, 1999), dykes can be considered as magma-filled cracks formed during tension associated with a transfersional setting.

In subduction zones, the displacement between the subducting slab and the overthrusted plate is nearly perfectly normal, and some obliqueness makes stress vectors suitable for space creation and pluton emplacement (Saint-Blanquat et al., 1998; Glazner, 1991). In this sense, an oblique convergence of the Caribbean plate with the continental margin of NW South America during Eocene time is considered as a likely tectonic scenario that permitted the emplacement of the shallower, cooler Parashi stock along with the simultaneous opening of E-W cracks that were substantially simultaneous filled by the dyke swarm in a transfensional setting. This setting created the numerous pull-apart-like basin that characterises northern South America in Colombia and Venezuela (Macellari, 1995; Montes et al., 2019).

The emplacement history of the adjacent Santa Marta batholith has been characterised as involving two phases. The first was a major emplacement phase (58-54 Ma) that has been associated with a regional synthetic extensional shear band, related to a NE-trending arc parallel shear plane that was accompanied with shortening (Salazar *et al.*, 2016). The second phase, occurring around 50 Ma, was contemporaneous with the emplacement of the Parashi stock (Fig. 9) that recorded the reactivation of the extensional structures. Together, these phases record the switch towards more oblique convergence of the tectonic plates in the northwestern margin of South America.

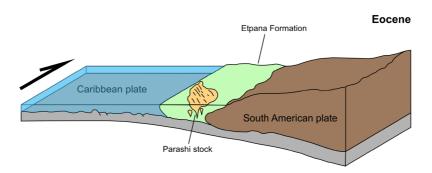


FIG. 9. By the Eocene, the intrusion of the Parashi stock was prompted by a transtensional tectonic regime resulting from the oblique convergence of the Caribbean plate with the continental margin of NW South America. During the final stages of the magma crystallisation, the pluton was able to deform by brittle fracturing, and the dykes intruded. The intrusion occurred when the overriding continental South American plate was deforming mostly by extension (modified from Weber *et al.*, 2010 and Cardona *et al.*, 2014).

6. Conclusions

Low-field AMS and AAR data, micro- and macrostructural evidence and barometric calculations of the Parashi stock emplacement and the related dyke swarm are all consistent with a shallow crustal emplacement (e.g., <10 km) in a transtensional tectonic setting related to the oblique convergence of the Caribbean and South American plates. This tectonic setting is similar to that proposed for another Eocene arc pluton, the Santa Marta batholith, although the latter records a net component of dextral shearing that was not recorded in the fabric of the Parashi stock. Although both plutons share emplacement mechanisms associated with the oblique convergence of the Caribbean plate, the Parashi stock may have been emplaced in a more stable and cooler crust and may be related to the interruption of the subduction in the Eocene. The magnetic foliation of the stock tends to show a concentric organization attributed to the differential upward movement of the magma. The dykes, in turn, were emplaced in the final stages of the magma crystallisation when the pluton was able to deform by brittle fracturing, which is consistent with the crystallisation ages determined for the stock and dykes.

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